



# Joint 21cm Parameter Constraints with Emulators

Nick Kern  
UC Berkeley

with Adrian Liu, Aaron Parsons, Andrei Mesinger and Bradley Greig

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Image: Alvarez et al. 2009

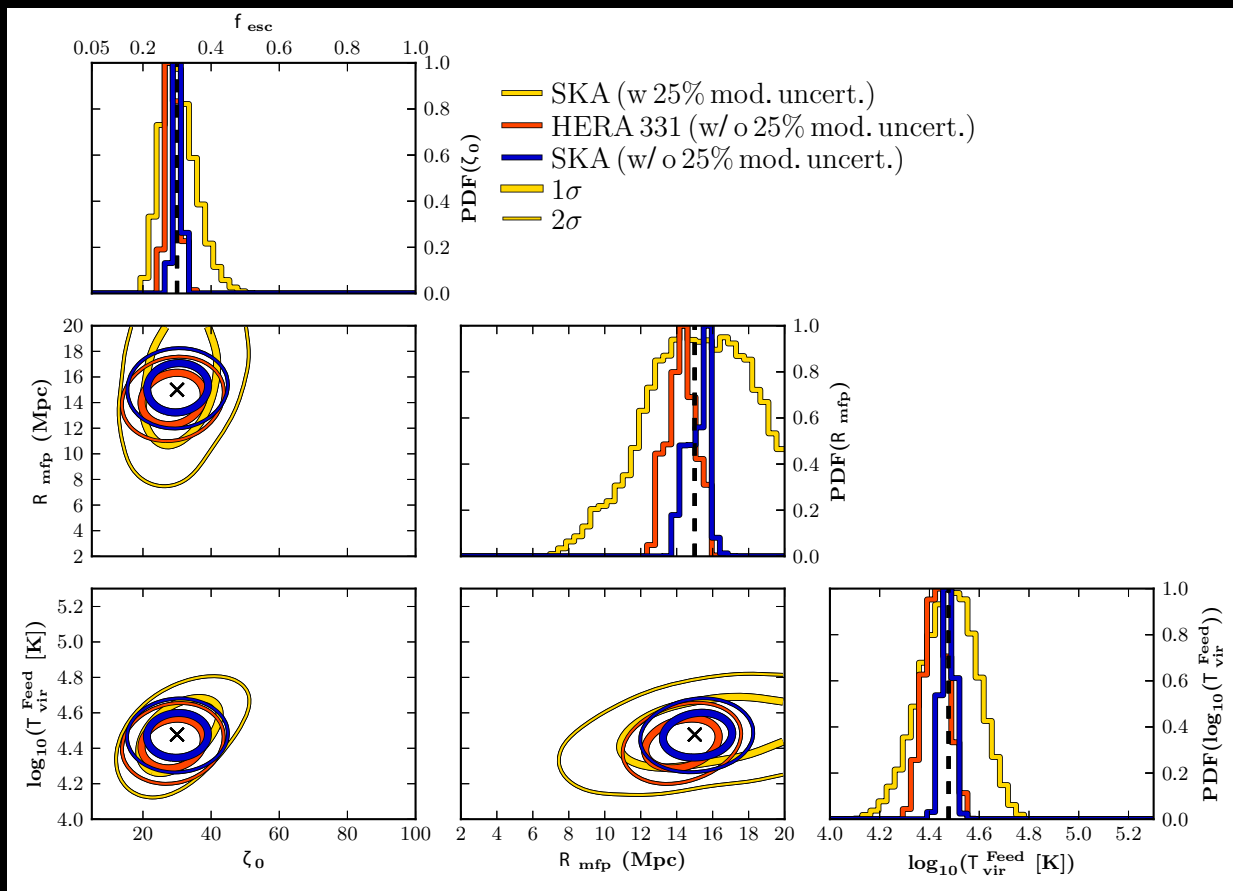
How do we interpret a 21cm power spectrum detection?

# How do we interpret a 21 cm power spectrum detection?

Start with a three parameter model for reionization

$R_{\text{mfp}}$

$T_{\text{vir}}$



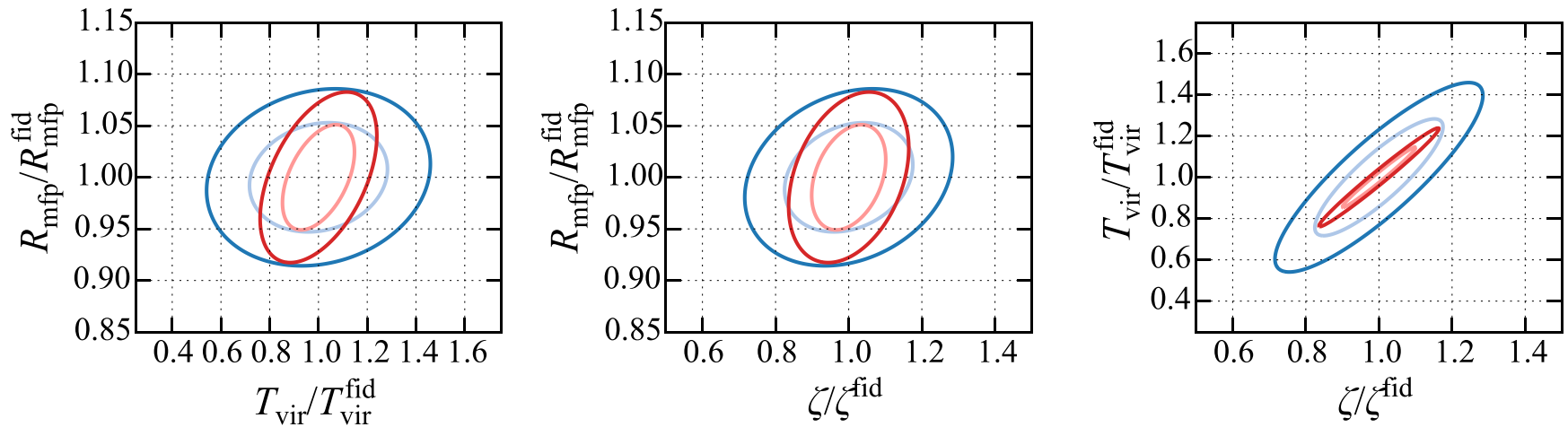
$\zeta$

$R_{\text{mfp}}$

Greig et al. 2016

# How do we interpret a 21cm power spectrum detection?

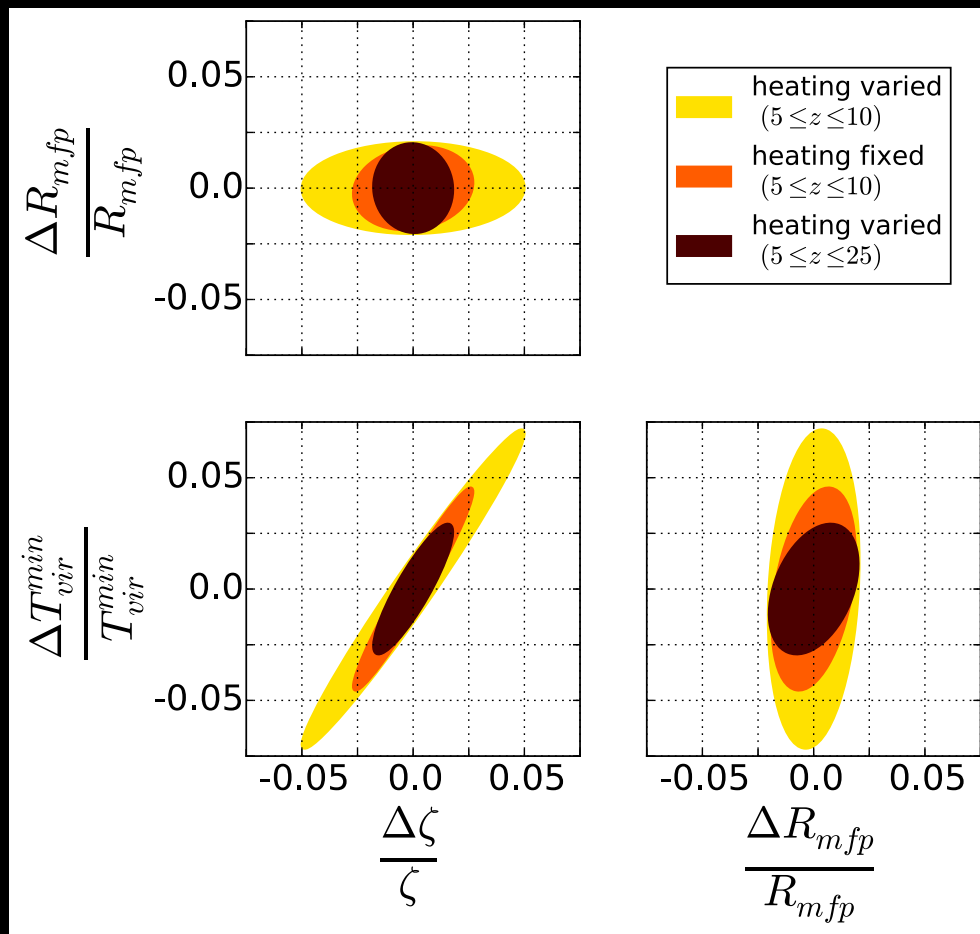
Cannot ignore cosmology



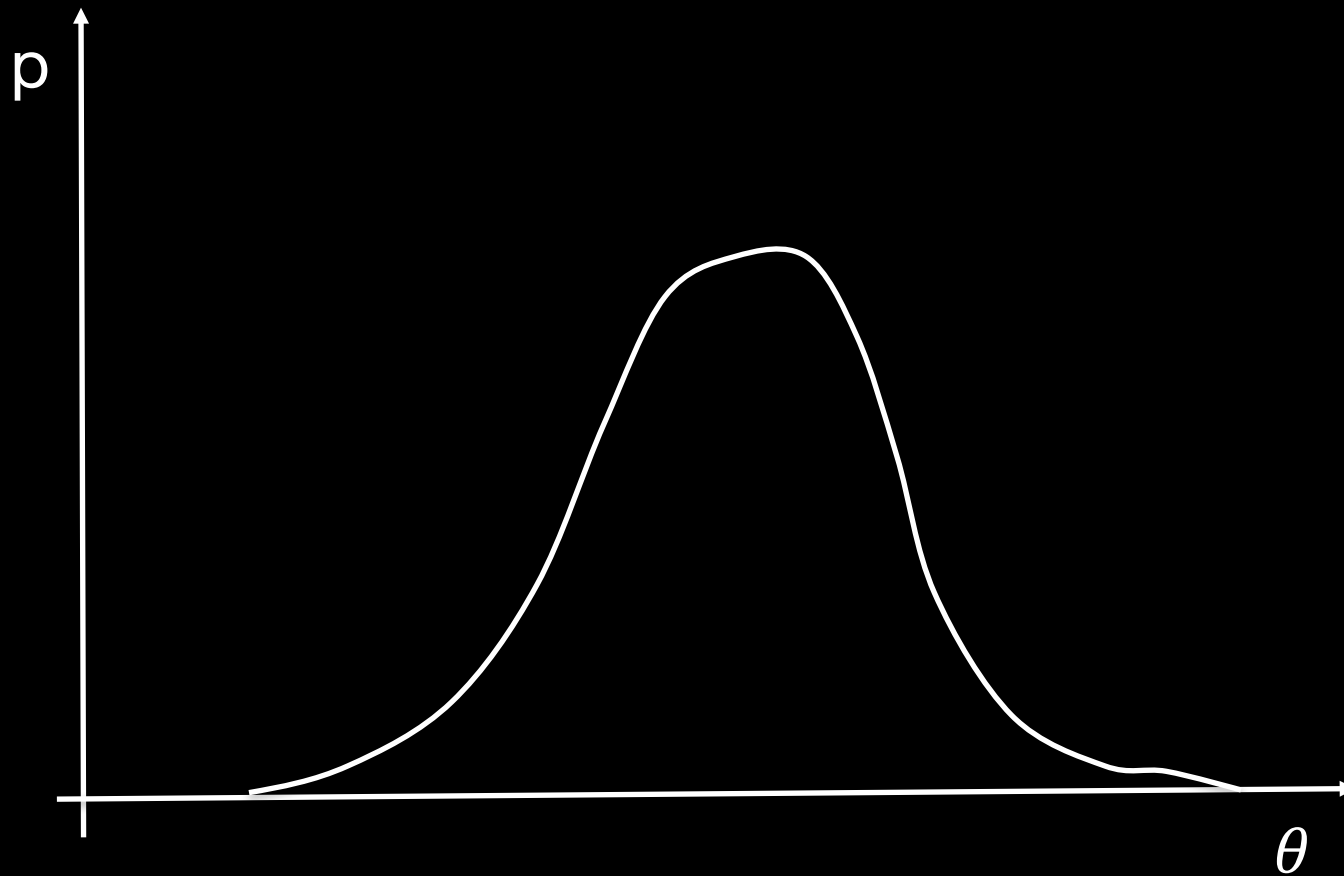
Liu et al. 2016

# How do we interpret a 21cm power spectrum detection?

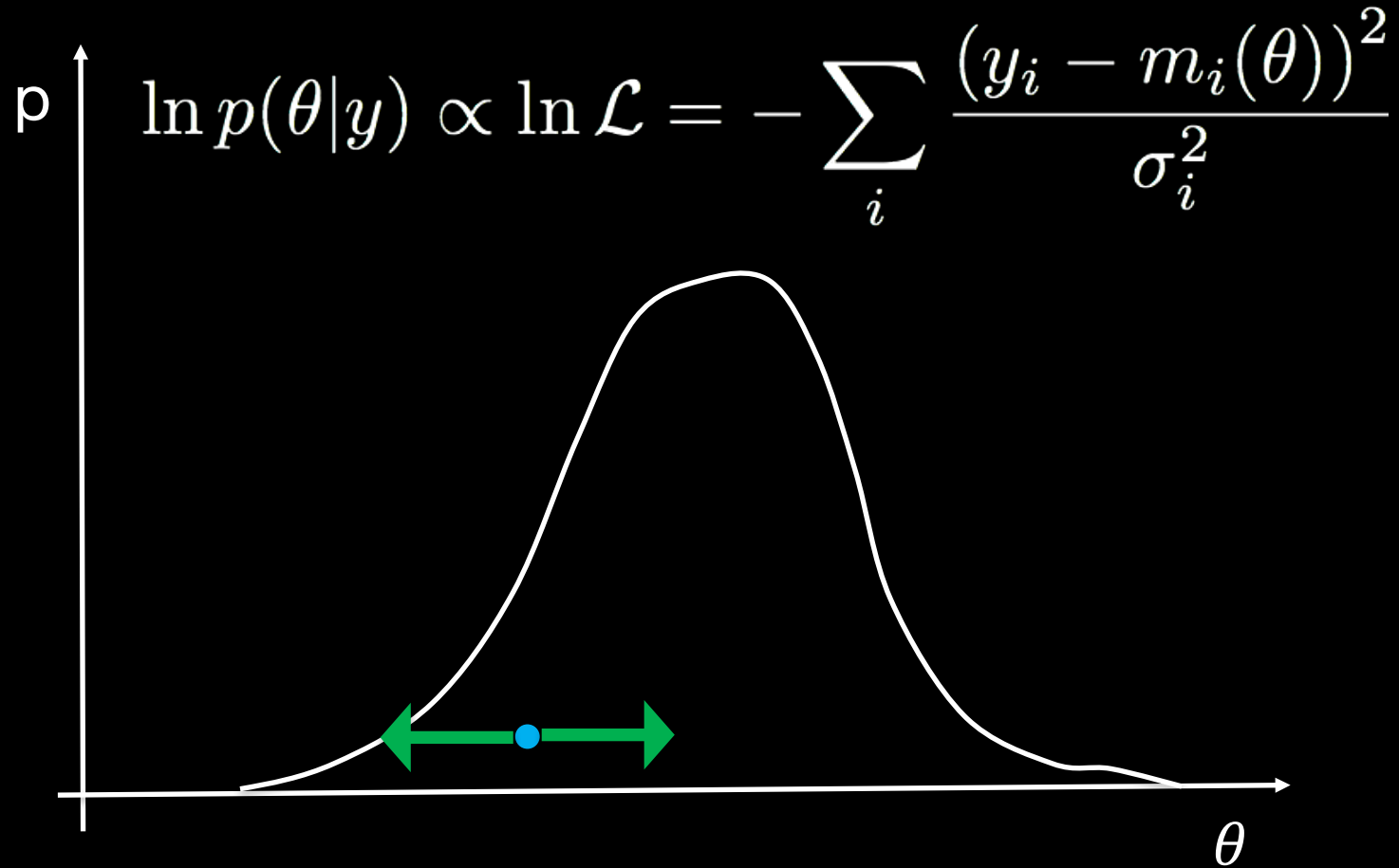
Should not ignore IGM heating



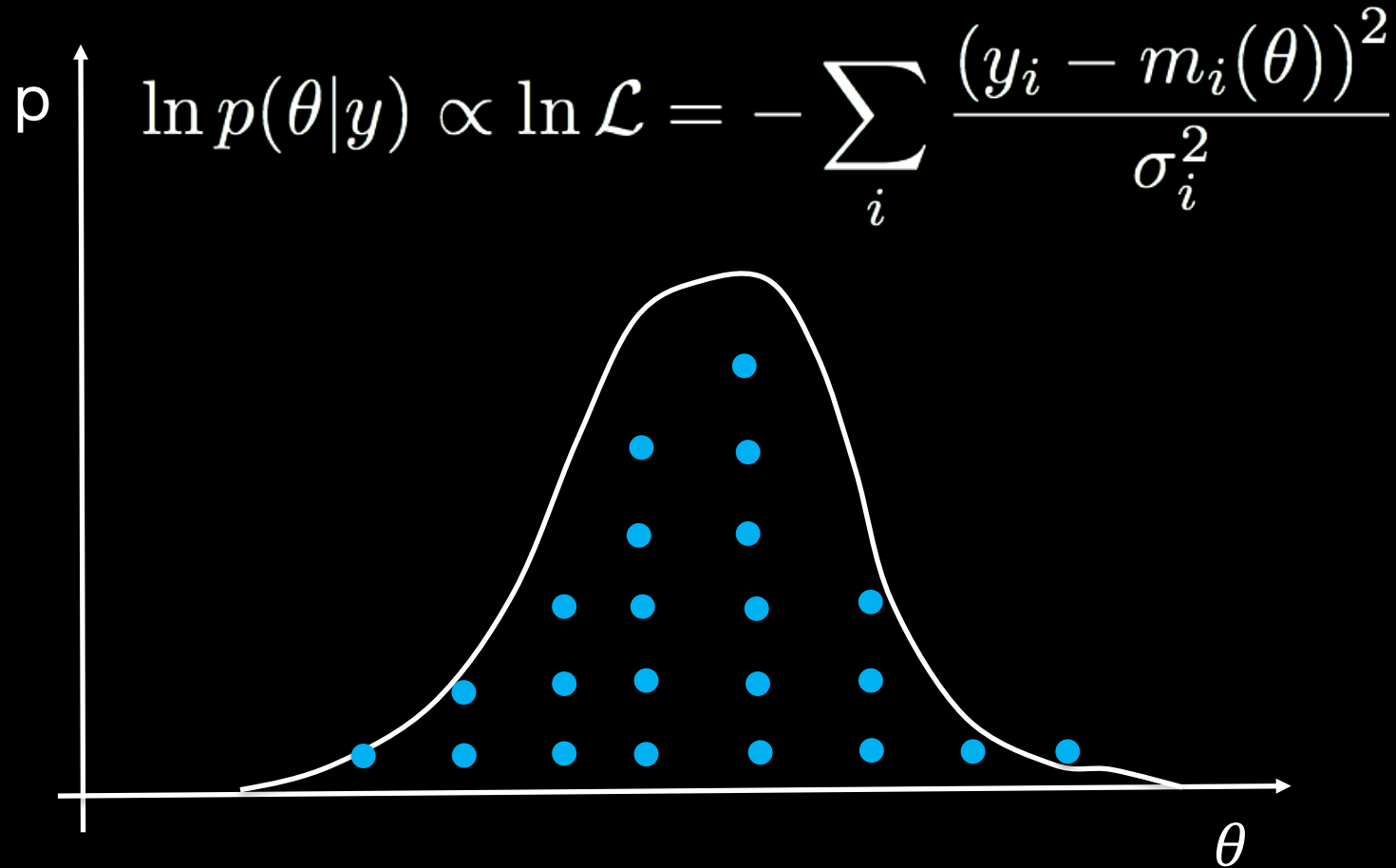
# Emulators Approximate the Likelihood



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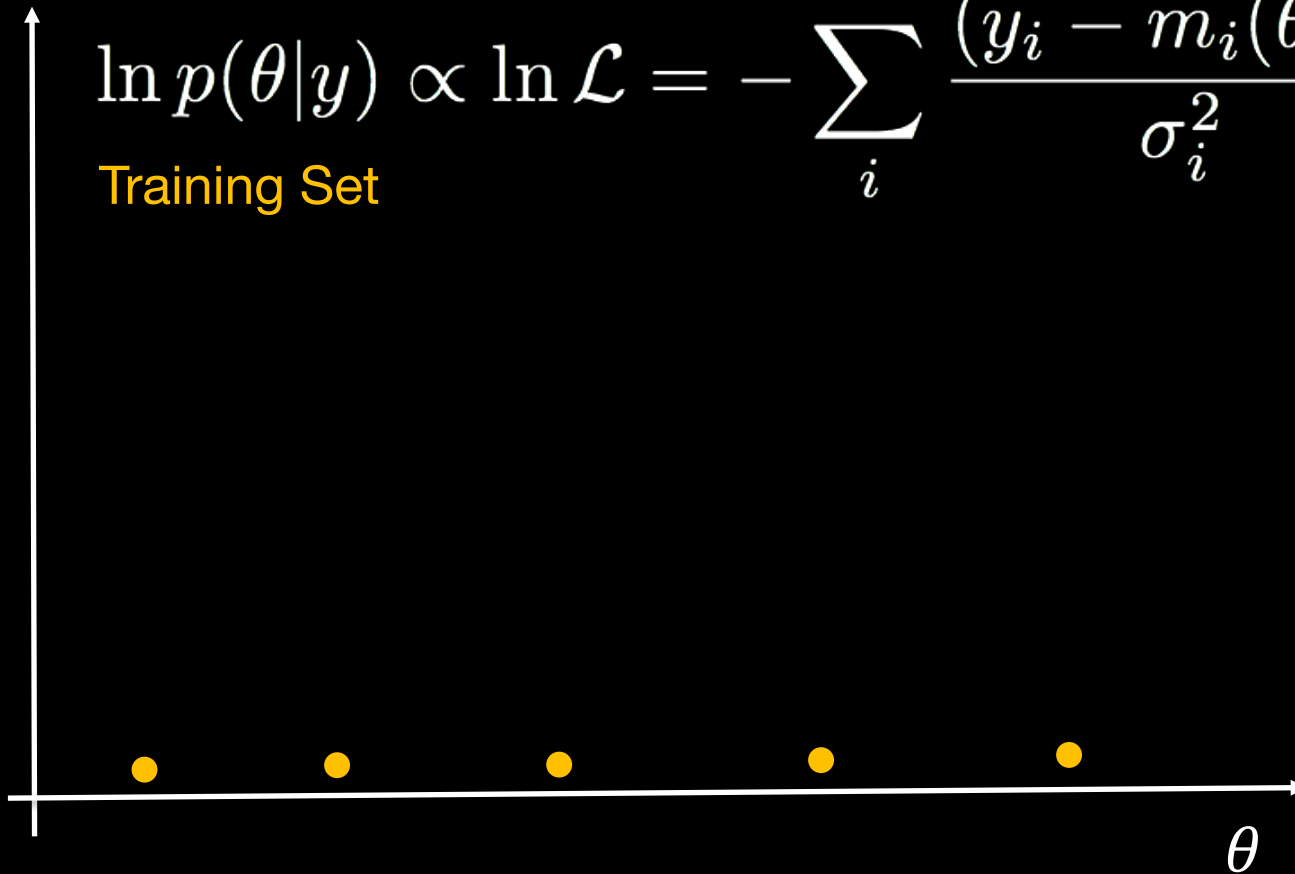




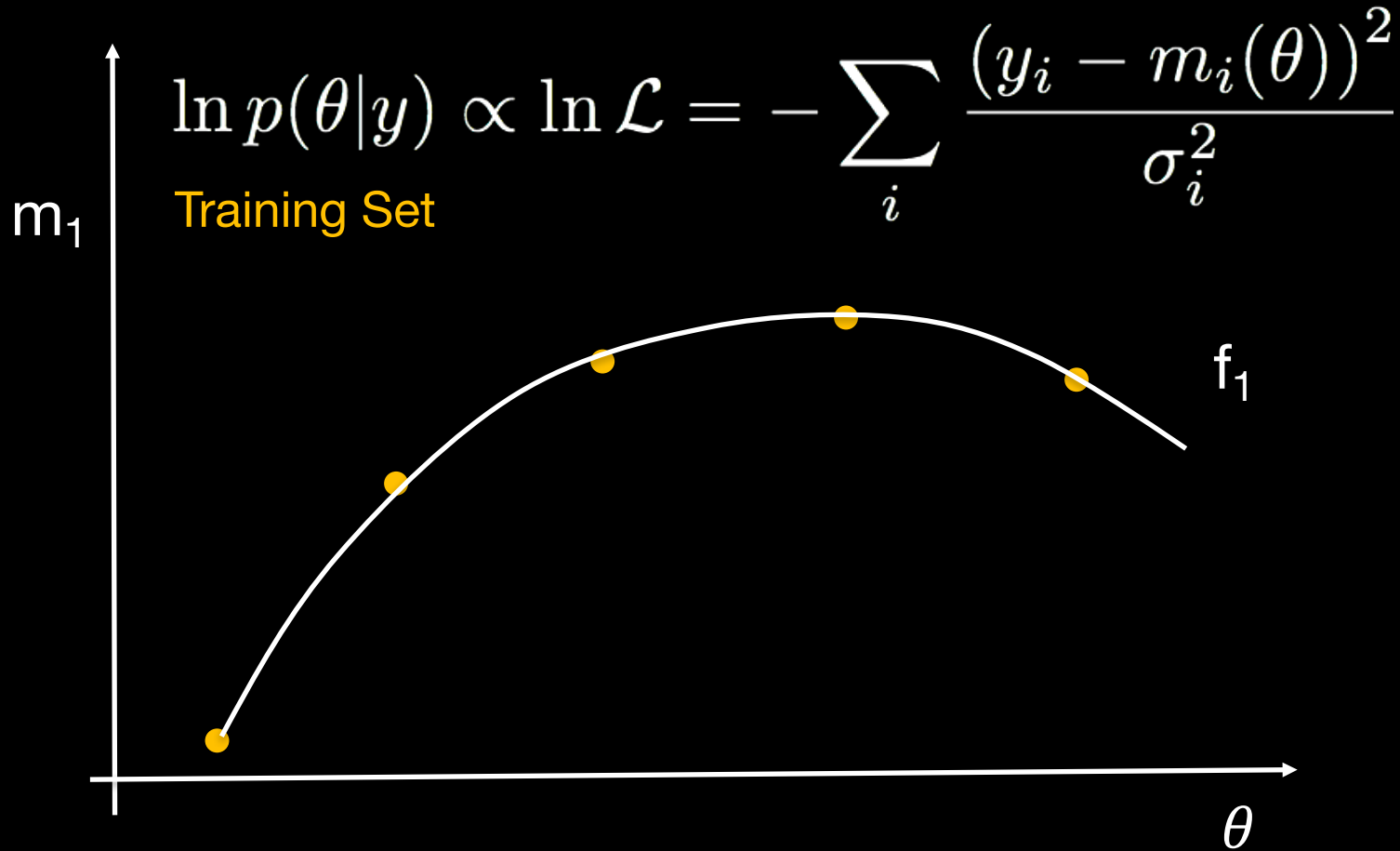
# Emulators Approximate the Likelihood

$$\ln p(\theta|y) \propto \ln \mathcal{L} = - \sum_i \frac{(y_i - m_i(\theta))^2}{\sigma_i^2}$$

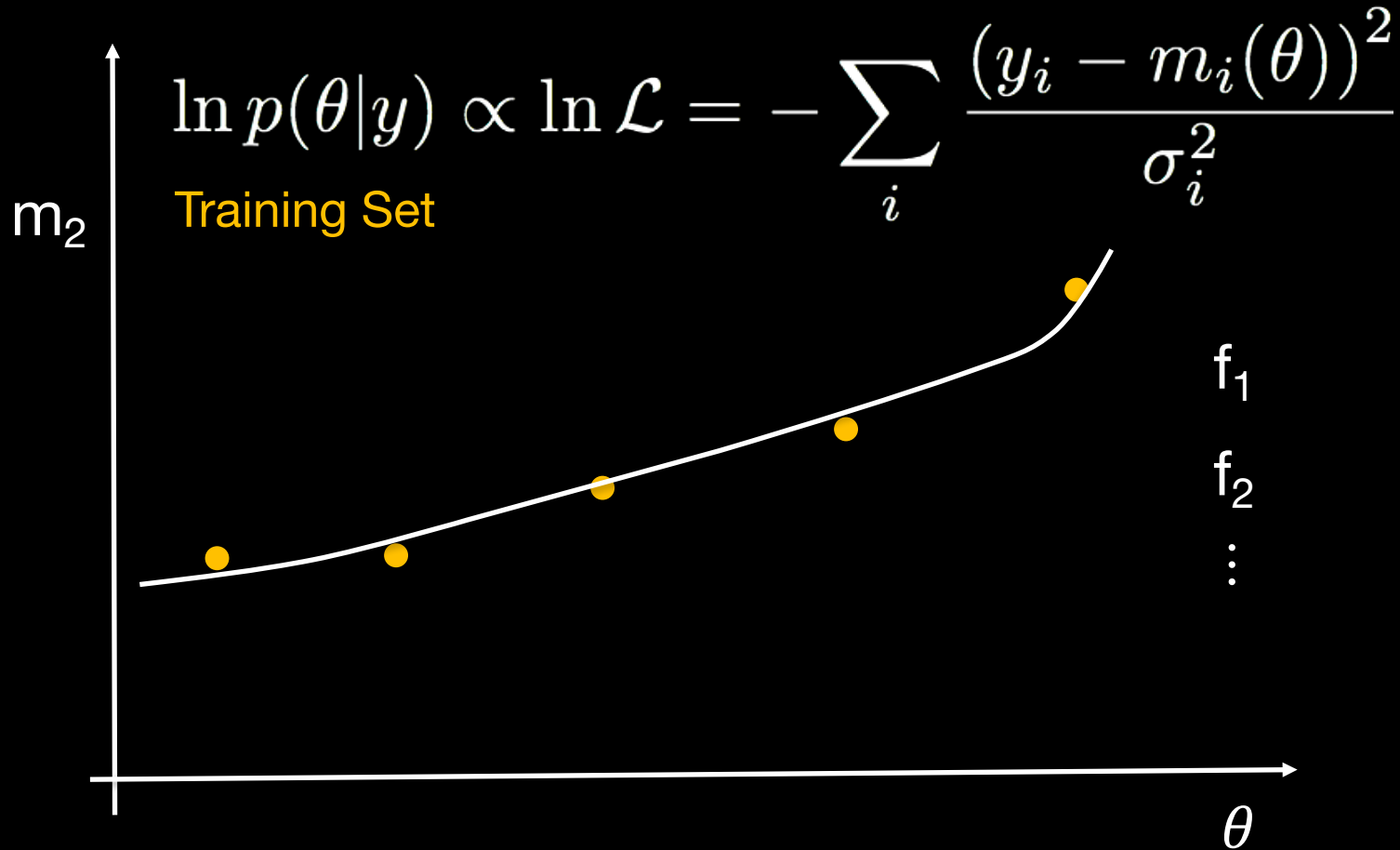
Training Set



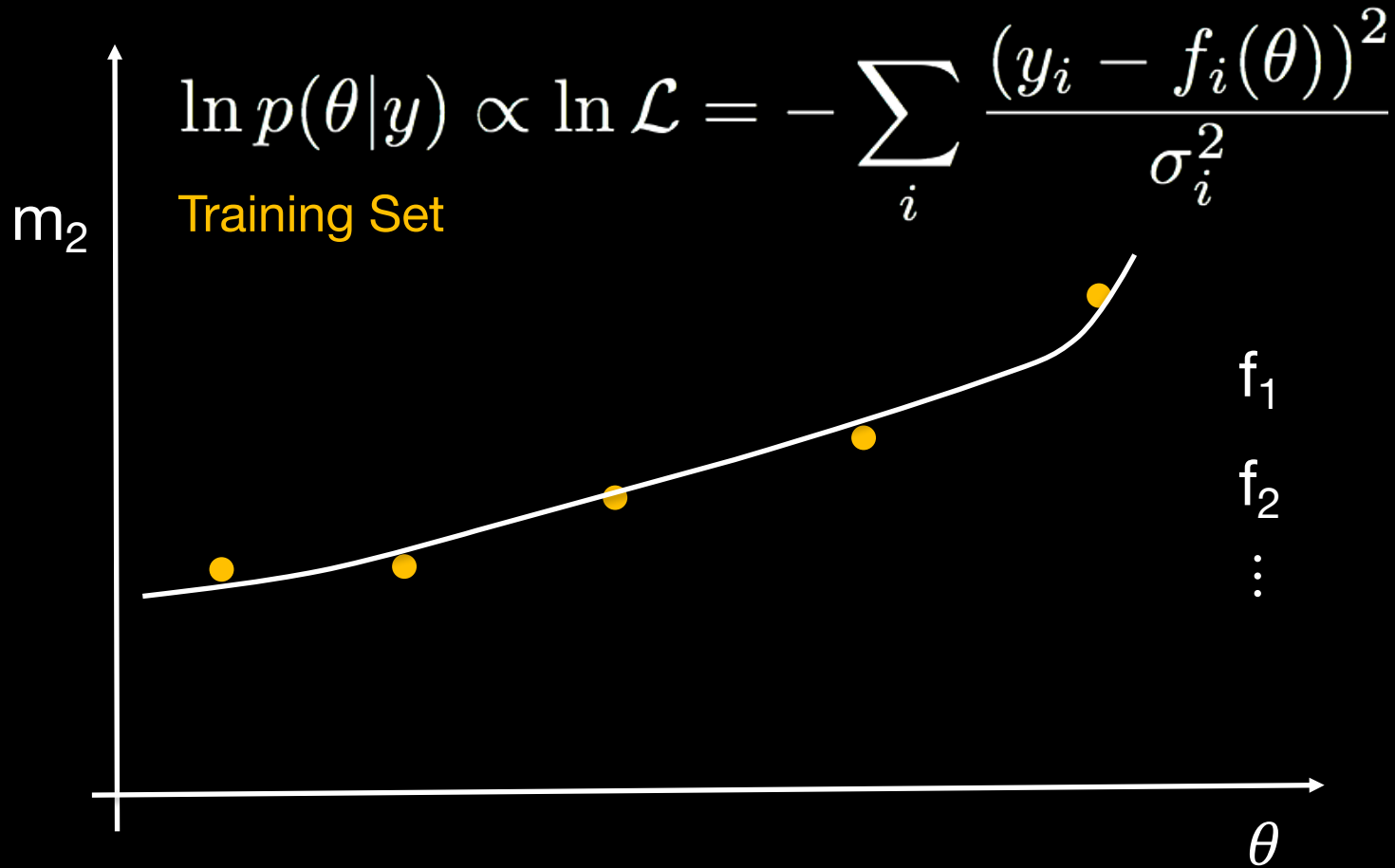
# Emulators Approximate the Likelihood



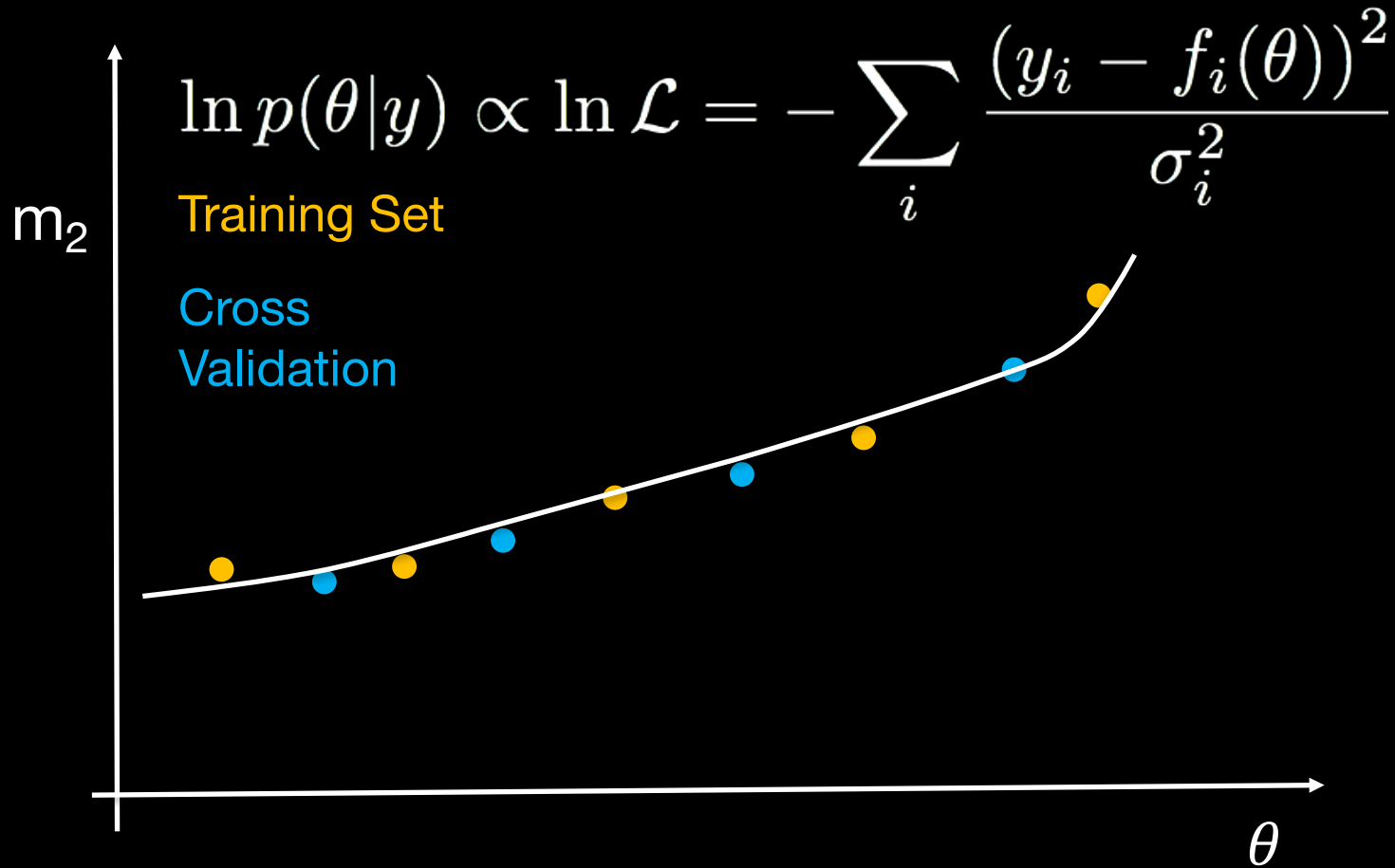
# Emulators Approximate the Likelihood



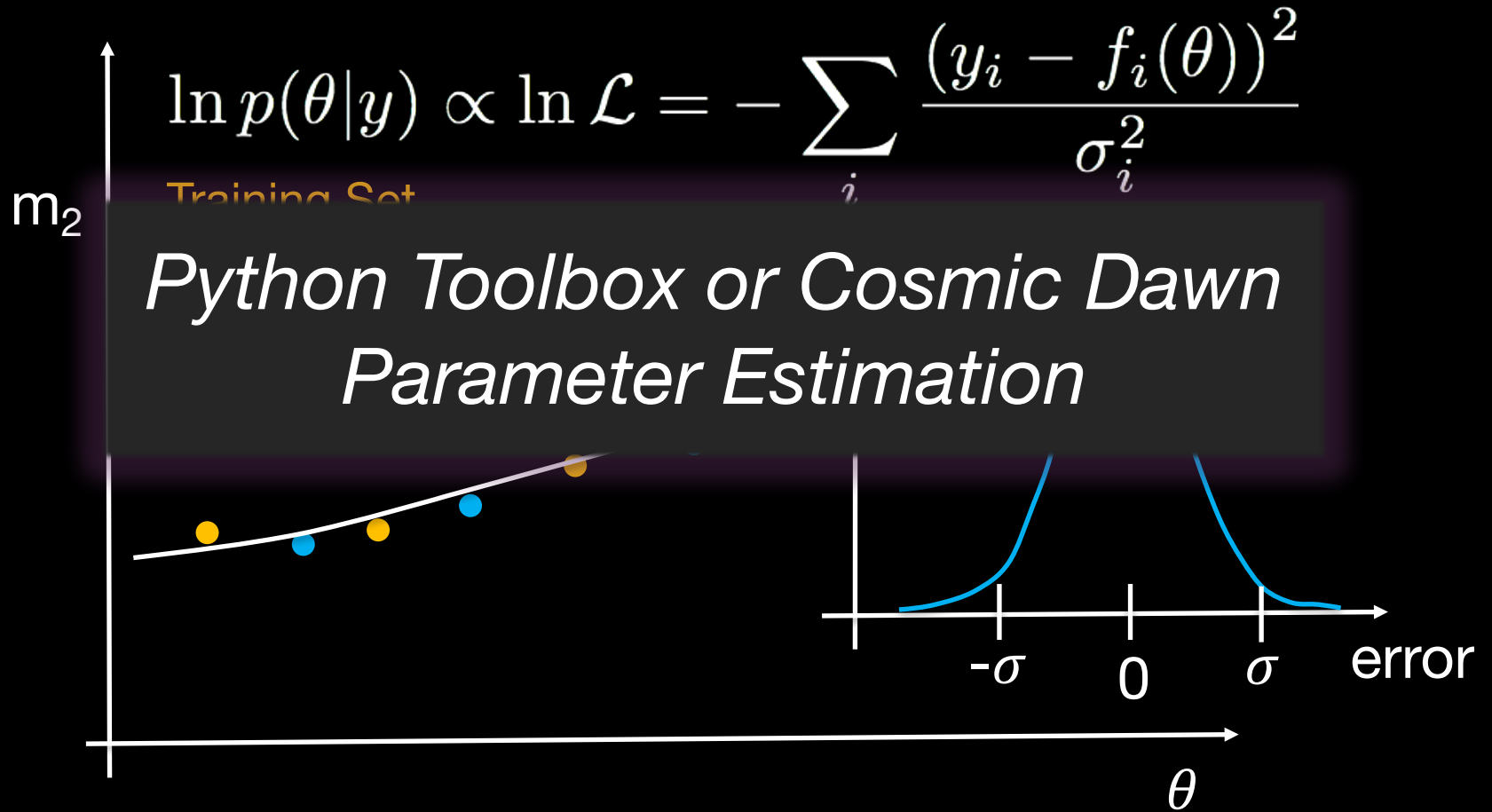
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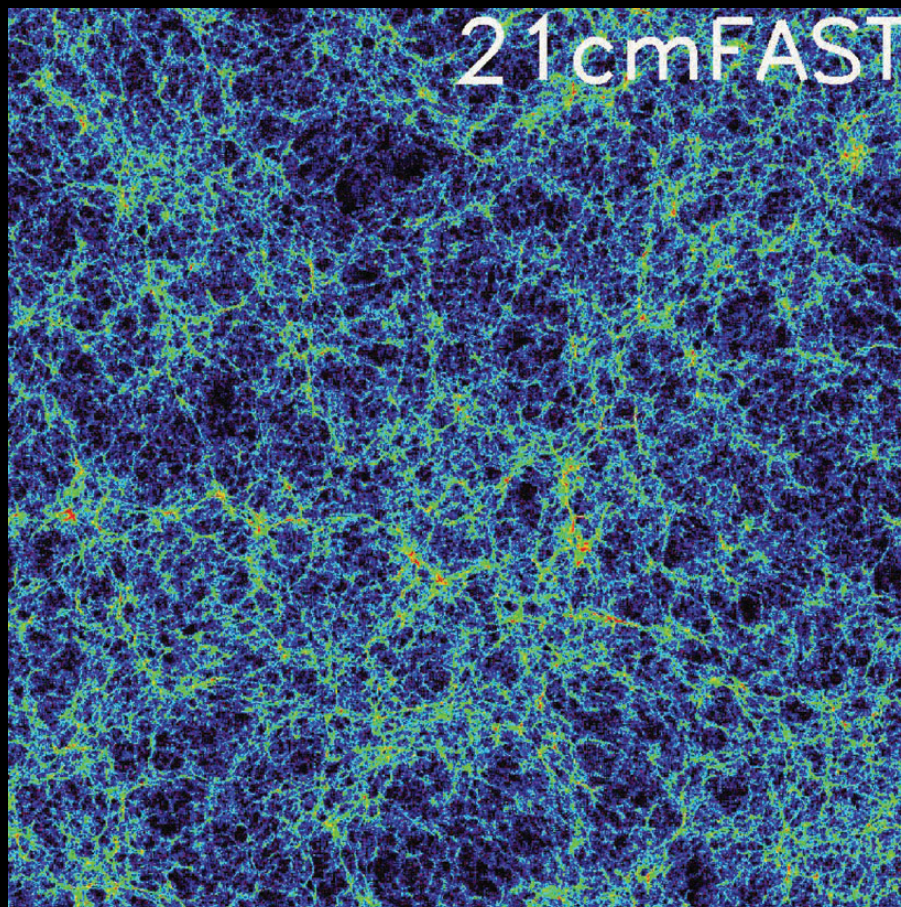


# Emulators Approximate the Likelihood



# HERA Parameter Constraint Forecast via Emulation of *21cmFAST*

# Training on EoR Simulations



$$L_{\text{BOX}} = 400 \text{ cMpc}$$

$$L_{\text{CELL}} = 2 \text{ cMpc}$$

Mesinger et al. 2011



# Parameters of Interest

## Astrophysics

- Ionization Efficiency:  $\zeta$
- Mean-free path of UV photons:  $R_{\text{MFP}}$
- Min. Virial Temp. of SF Halos:  $T_{\text{VIR}}$
- X-ray Efficiency:  $f_{\text{X}}$
- X-ray Spectral Index:  $\alpha_{\text{X}}$
- X-ray Cutoff Frequency:  $\nu_{\text{min}}$

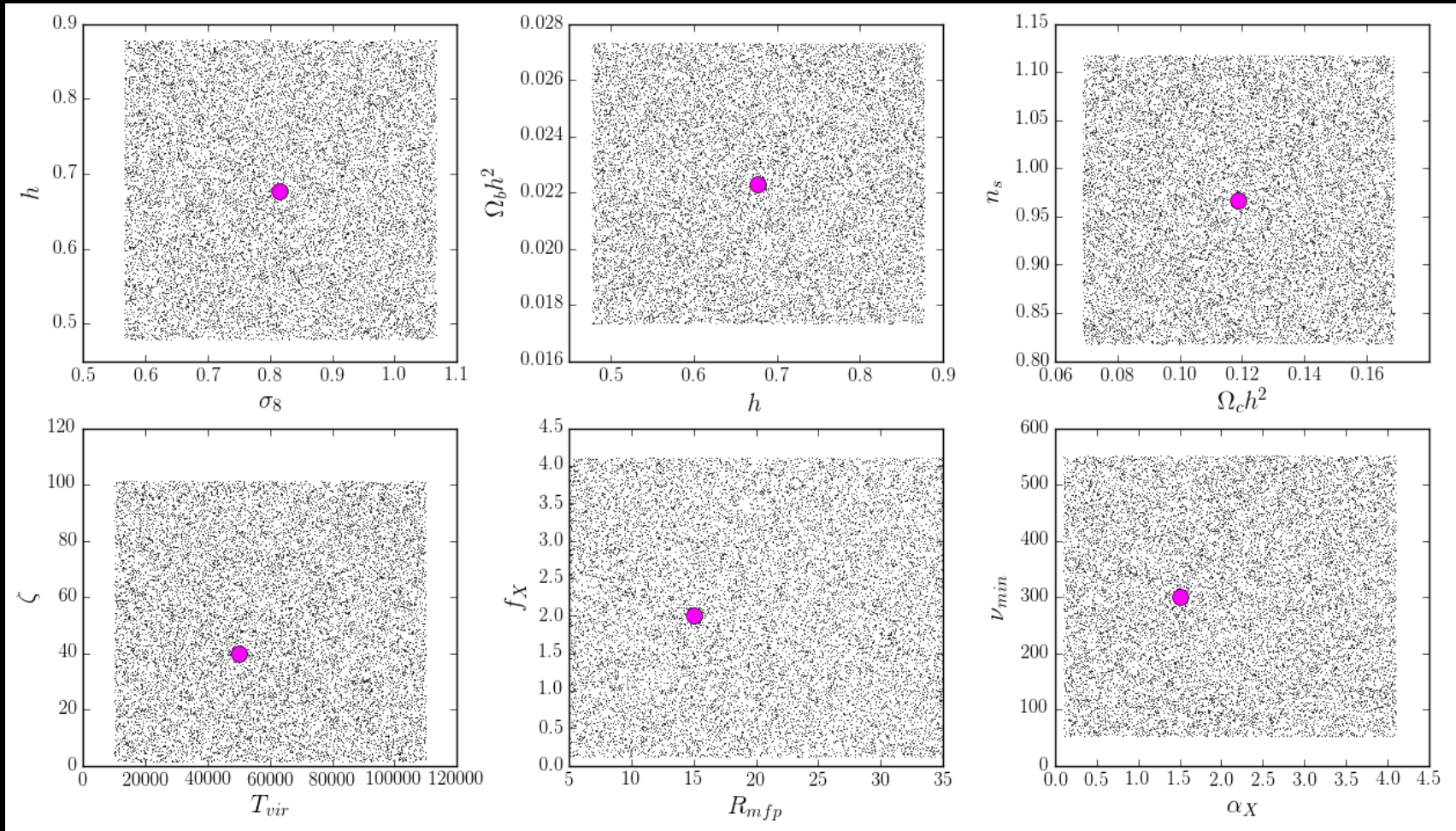
flat priors

## Cosmology

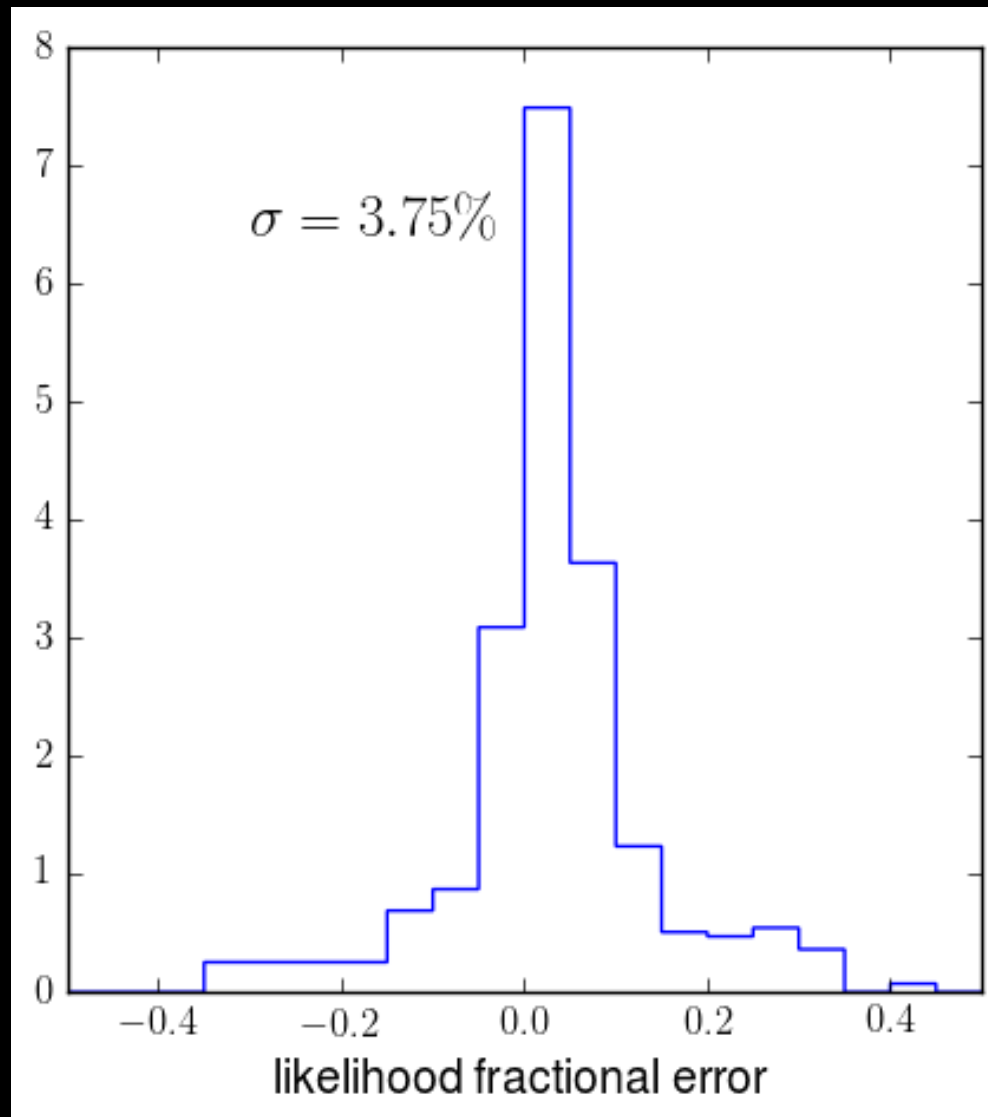
- Clustering Amplitude:  $\sigma_8$
- Primordial PS Slope:  $n_s$
- CDM Fraction:  $\Omega_{\text{c}}h^2$
- Baryon Fraction:  $\Omega_{\text{b}}h^2$
- Hubble Constant:  $H_0$

planck priors

# Emulation of Power Spectra at Percent-Level Accuracy

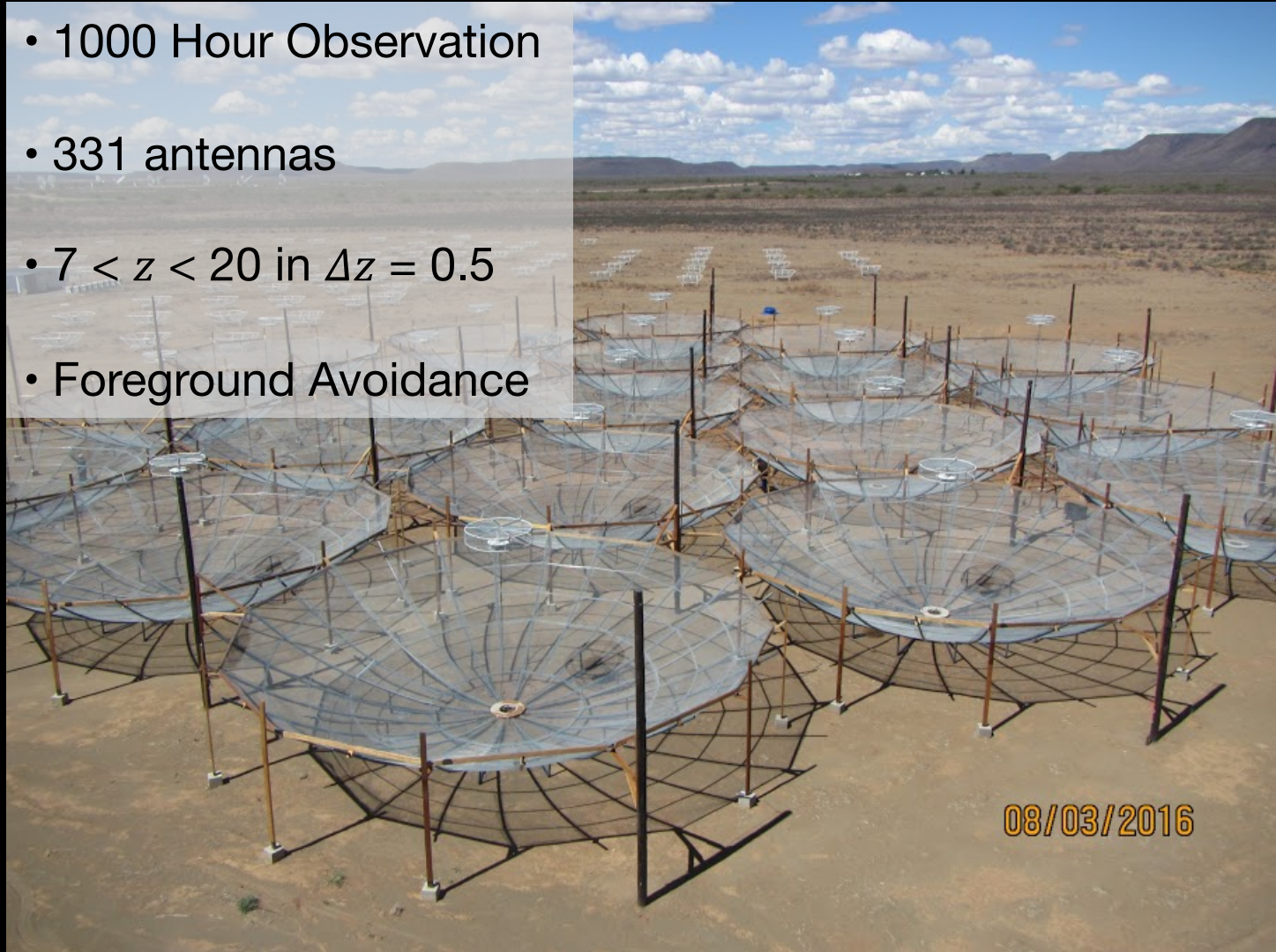


# Emulation of Power Spectra at Percent-Level Accuracy

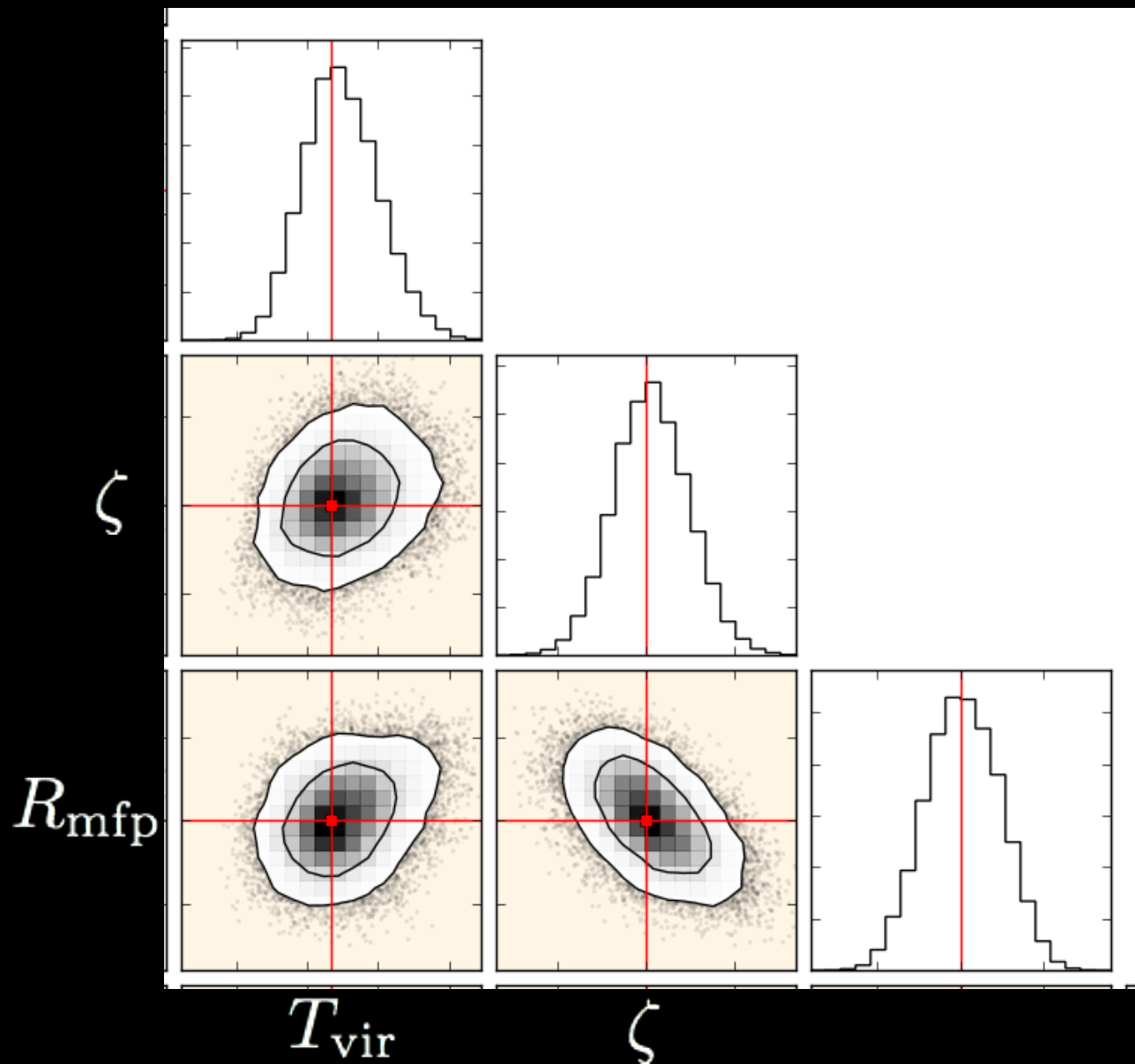


# HERA331 Power Spectrum Mock Observation

- 1000 Hour Observation
- 331 antennas
- $7 < z < 20$  in  $\Delta z = 0.5$
- Foreground Avoidance

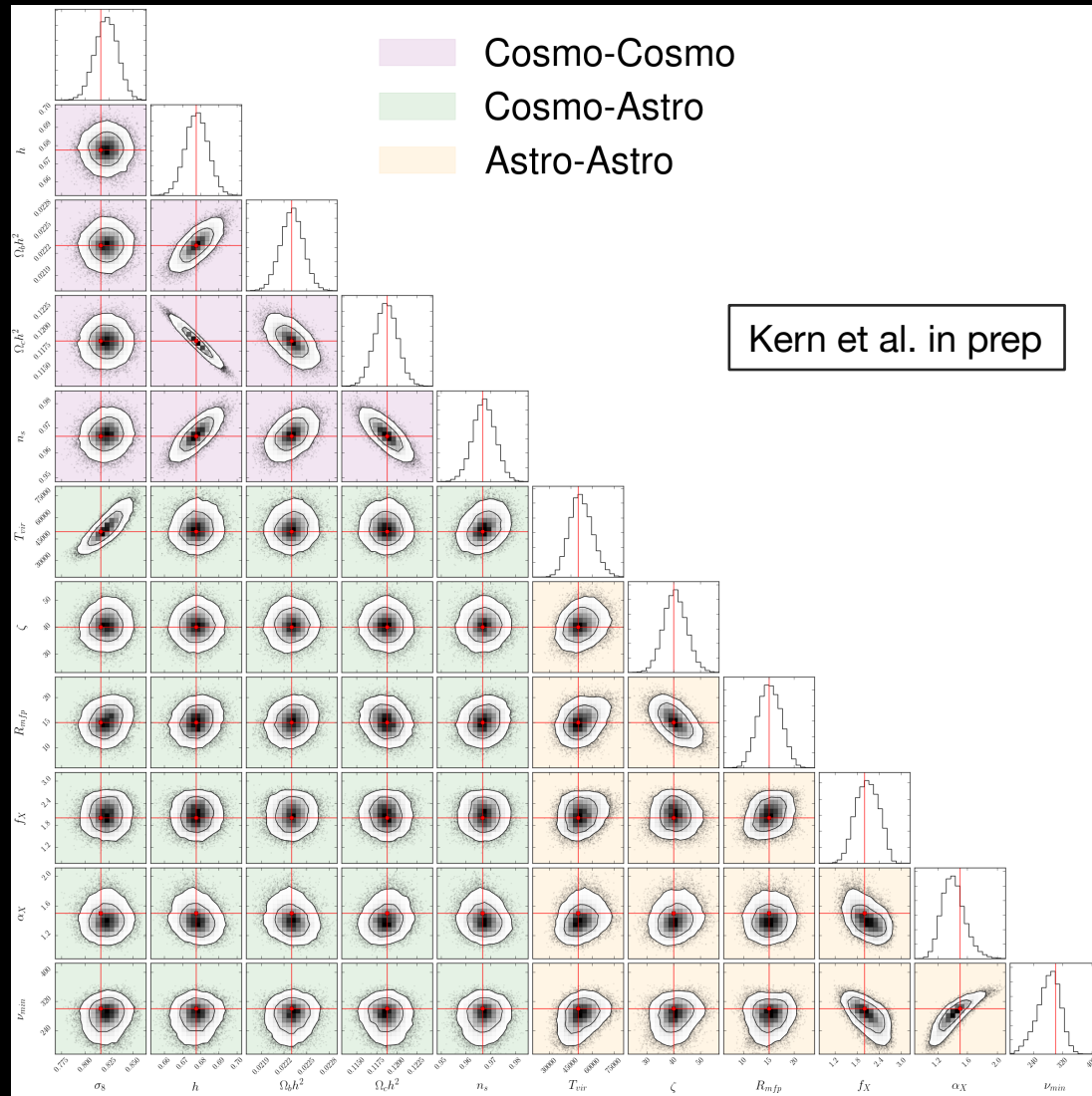


# HERA Joint Posterior Distribution

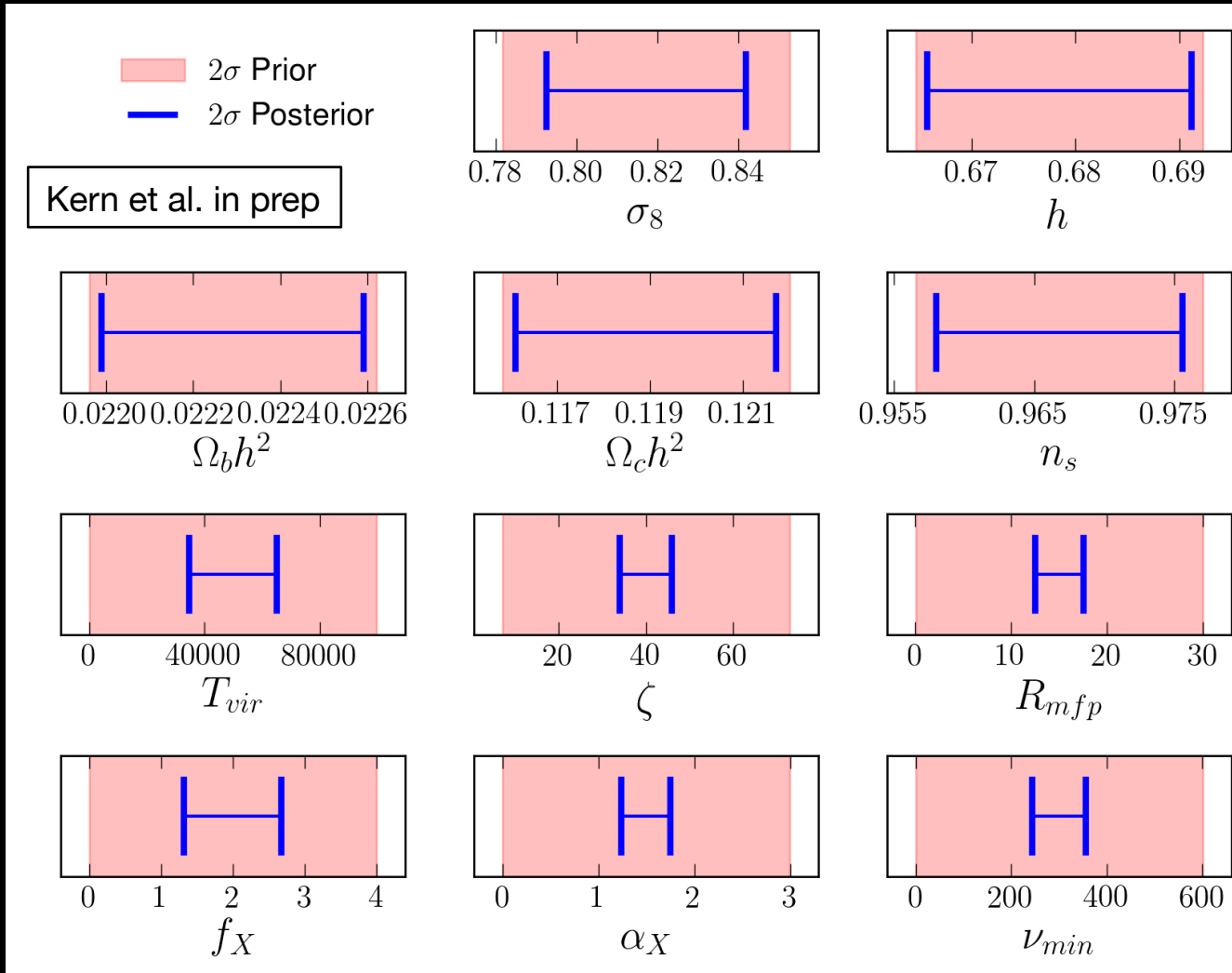


# HERA Joint Posterior Distribution

$\sigma_8$   
 $h$   
 $\Omega_b h^2$   
 $\Omega_c h^2$   
 $n_s$   
 $T_{\text{vir}}$   
 $\zeta$   
 $R_{\text{mfp}}$   
 $f_X$   
 $\alpha_X$   
 $\nu_{\text{min}}$



# HERA Marginalized Posterior Widths

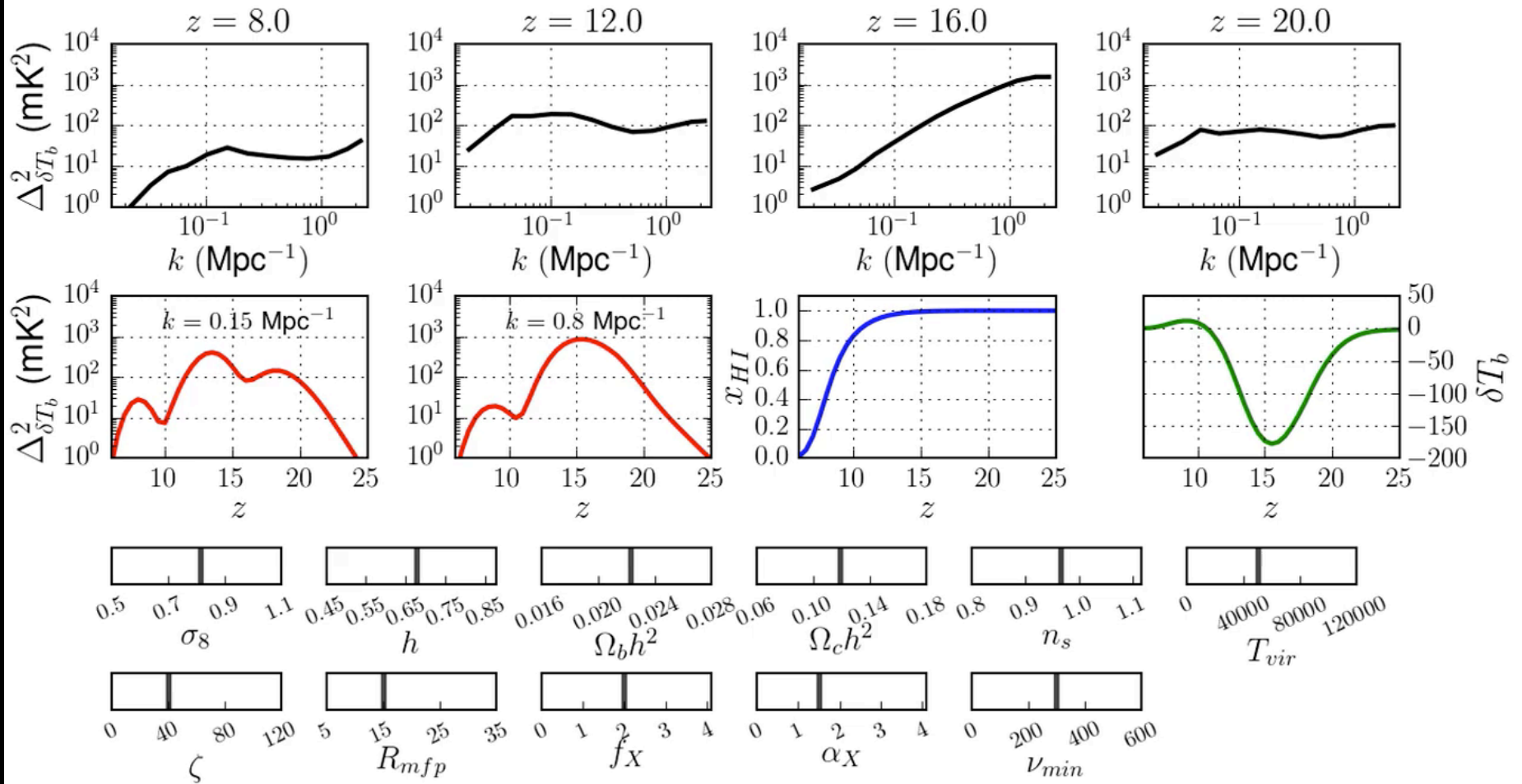


# Take-Away Points

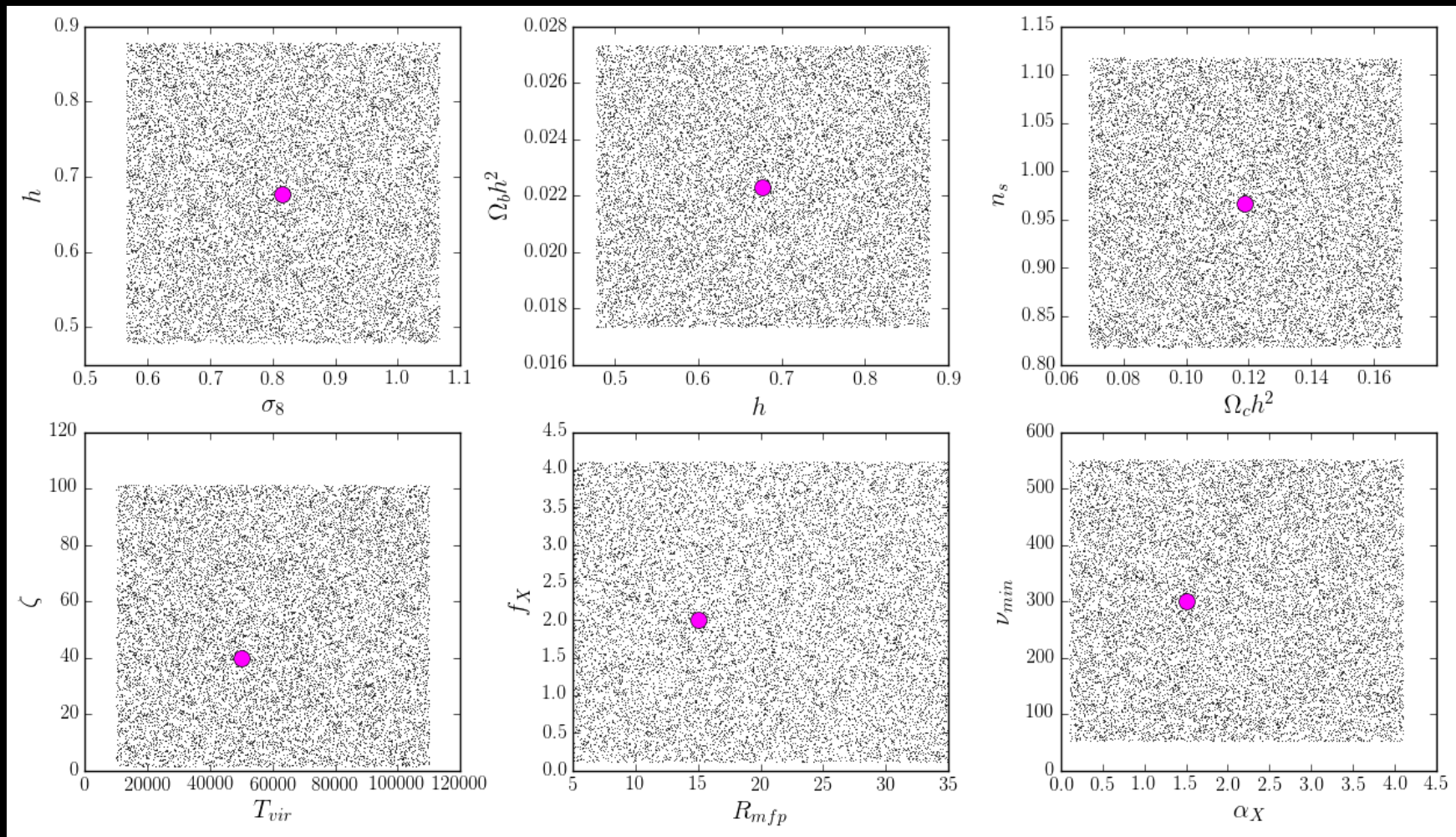
- 1 Emulators allow us to use more accurate simulations for full MCMC sampling of the posterior distribution
- 2 Emulators are generalizable to *any* simulation
- 3 HERA is predicted to put significant constraints on astrophysical parameters governing reionization and the heating epoch
- 4 The *Python Toolbox for Cosmic Dawn Parameter Estimation* will soon be a publicly available package!



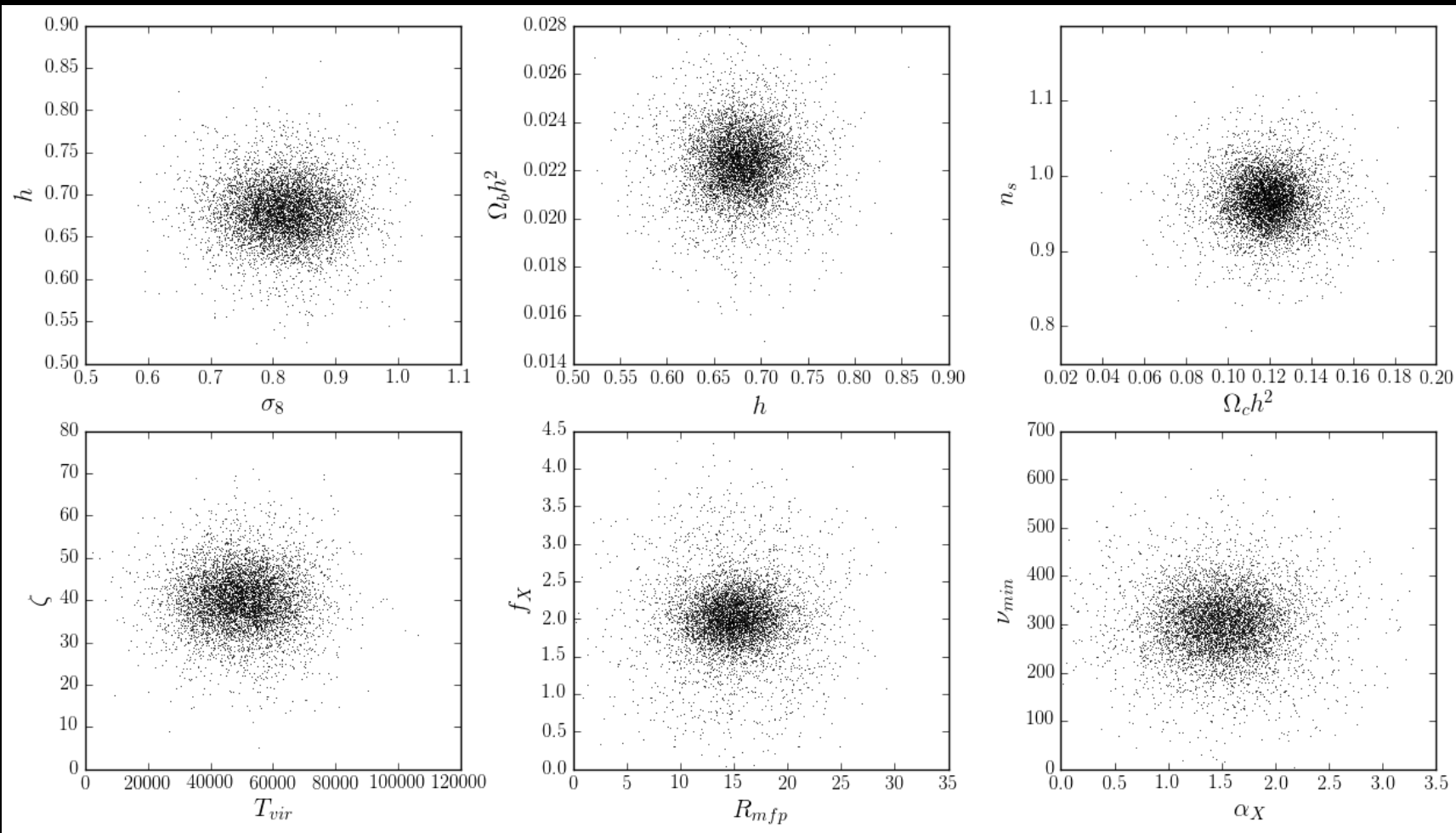
# Gaining Intuition



# Emulation of Power Spectra at Percent-Level Accuracy



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